

Proposed Laboratory Measurement of the Gravitational Repulsion Predicted by Quantum Celestial Mechanics (QCM)

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Quantum Celestial Mechanics (QCM) predicts the quantization of the orbital angular momentum per unit mass for bodies orbiting a central mass in response to attractive and repulsive gravitational accelerations. Applications to the Solar System, multi-planet exosystems, and to the Pluto system of 5 moons suggest its validity. A laboratory experiment to check this constraint is proposed.

1 Introduction

The gravitational constant G has now been measured by several new techniques, including a dynamic measurement by resonating beams [1] and a simple pendulum laser interferometer [2]. Both methods as well as Advanced LIGO and other gravitational sensors could also measure the repulsive gravitational acceleration predicted by the quantization of angular momentum per unit mass constraint [3] of Quantum Celestial Mechanics (QCM).

Although the Pluto system with its 5 satellites has already been a definitive test of this constraint [4], and its successful applications to the Solar System and numerous multi-planet exosystems have been achieved [5], an Earth-bound laboratory measurement confirmation is preferred.

According to QCM, which is derived from the general relativistic Hamilton-Jacobi equation, the quantization of orbital angular momentum L per unit mass μ constraint of the orbiting body, with quantization integer m , depends upon the total angular momentum L_T for the system of total mass M_T as

$$L/\mu = m L_T / M_T. \quad (1)$$

Recall that all orbits are equilibrium orbits for Newtonian gravitation for a central mass M and orbit distance r because the radial acceleration

$$\ddot{r} = -\frac{GM}{r^2} + \frac{L^2}{\mu^2 r^3}. \quad (2)$$

But for QCM, the subset of allowed equilibrium orbits are the ones that obey

$$\ddot{r} = -\frac{GM}{r^2} + \frac{m(m+1)L_T^2}{M_T^2 r^3} \quad (3)$$

for circular orbits. Therefore, a very small radial displacement from the equilibrium radius r_{eq} of orbit results in an acceleration in the opposite direction.

2 Lab experiment parameters

In order to mimic a Keplerian circular orbit, one would place an ideal rotating metal cylinder of mass M and radius R at

a distance r from the gravitational detector. A simple estimation of the parameters for a laboratory scale measurement is made by assuming that the detector is essentially a point mass M_d responding instead of an extended geometrical object. Therefore,

$$r_{eq} = \frac{m(m+1)L_T^2}{GM M_T^2} \approx \frac{m(m+1)R^4 \omega^2 M}{4G(M+M_d)^2}. \quad (4)$$

Inserting some reasonable values: $M = 5$ kg, $R = 5$ cm, $M_d = 2$ kg, and $m = 1$, the first equilibrium radius will be at $r_{eq} \approx 4781\omega^2$ metre. For $r_{eq} = 1$ metre, i.e. fit in a lab room,

$$\omega \approx 0.0145 \text{ rad/s} \approx 8.3 \text{ rot/hr}. \quad (5)$$

By varying the rotation rate ω of the cylinder one can sweep back and forth through several equilibrium radii for $m = 1, 2, 3, \dots$ to observe attractive and repulsive accelerations at $r_{eq} = 2r_0, 6r_0, 12r_0, \dots$ sensed by the detector, with rapidly decreasing interaction accelerations with increasing r_{eq} .

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